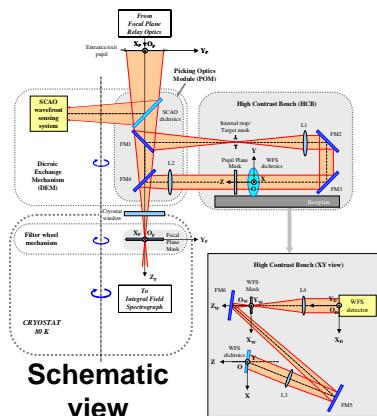


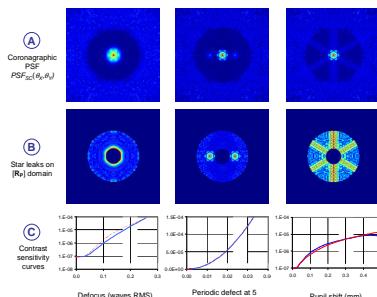
# Opto-mechanical design of a High Contrast Module (HCM) for HARMONI

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Schematic view



System budget

Type of perturbation	Amplitude	Sensitivity coefficient $\alpha$	Power $p$	Weight $w$	Contrast C
Image quality	0.6 (waves RMS)	-	-	-	8.7E-08
Nominal	0.06	-	-	-	
Tilt Alignment	0.05	5.7E-04	4	1	3.6E-09
Atmospheric dispersion	0.39	5.7E-04	4	1	1.3E-05
Focus	0.02	1.6E-02	4	1	2.5E-09
Astigmatism	0.05	4.5E-02	5	1	1.4E-08
Oscillatory Low frequency (< 5 cycles)	0.01	1.2E-02	2	0.5	6.2E-07
Mid frequency (5-20 cycles)	0.01	9.0E-02	2	0.5	4.5E-06
High frequency (> 20 cycles)	0.02	2.0E-02	2	0.5	4.0E-06
Sub-total RMS	<b>0.10</b>				<b>1.5E-05</b>
Pupil shift	x (mm)				
Lateral alignment	0.2	5.6E-05	2	1	2.3E-06
Total RMS					<b>1.5E-05</b>
Science requirement					<b>1.5E-05</b>

**Summary**

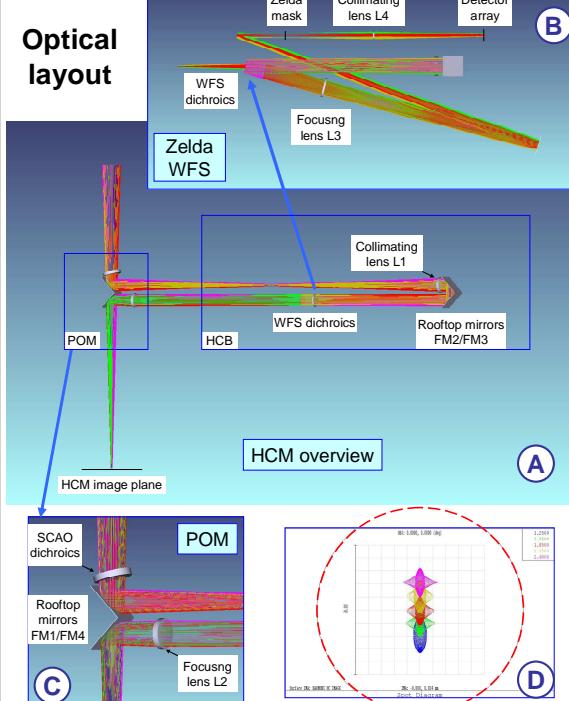
HARMONI is a first-light visible and near-IR integral field spectrograph of ESO's Extremely Large Telescope. A Single Conjugate Adaptive Optics sub-system will provide diffraction-limited spectral images in a Nyquist-sampled  $0.61 \times 0.86$  arcsec field of view with a  $R=3000\text{-}20000$  spectral resolution. The High Contrast Module (HCM) will add an essential high-contrast imaging capability for HARMONI to spectrally characterize young giant exoplanets and disks with flux ratio of  $1E\text{-}6$  at  $0.1\text{-}0.2''$  from their star. It uses an apodized pupil coronagraph to lower the intensity of the diffracted starlight and limit the dynamic range on the detector, and an internal wavefront sensor to calibrate non-common path aberrations.

## From science to technical requirements

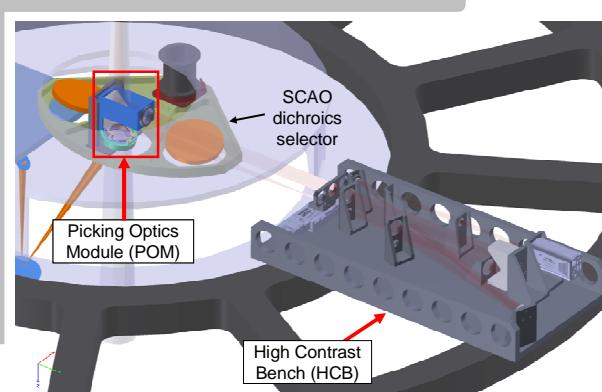
SCIENCE REQUIREMENTS	VALUES
Spectral range	Science: From 1.25 to 2.45 $\mu\text{m}$ Wavefront sensing: From 1.20 to 1.25 $\mu\text{m}$
Contrast (see definition below)	$\leq 1.5 \times 10^5$ at 200 milli-arcsec (mas) $> 50\%$ in full science spectral range
Throughput (including PPM apodization, see definition below)	860 $\times$ 610 mas sampled at 4 mas
FoV and FoV sampling	$> 300$ mas $< 75$ mas
Outer Working Angle (OWA)	$\lambda/100$ RMS in full science spectral range
Inner Working Angle (IWA)	
WFE measurement accuracy (performed by built-in WFS)	

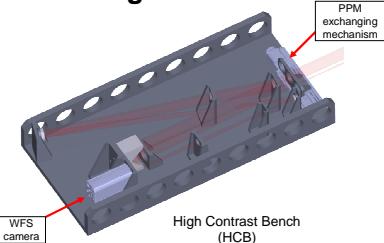
TECHNICAL REQUIREMENTS	VALUES
Optics transmission (excluding PPM)	$> 85\%$ in full science spectral range
Image position and stability (at FPRO focal plane)	20 and 2 $\mu\text{m}$ respectively along $X_F$ and $Y_F$ axes 200 and 20 $\mu\text{m}$ respectively along $Z_F$ axis
Image quality	Global (tilt and focus removed) Low spatial frequency errors (< 5 cycles) Mid spatial frequency errors (5-20 cycles) High spatial frequency errors (> 20 cycles)
Exit pupil position and stability	$< 32$ nm RMS (goal 17 nm) $< 17$ nm RMS (goal 9 nm) $< 20$ nm RMS (goal 10 nm) $< 18$ nm RMS (goal 10 nm) $< 0.25$ and $0.1\%$ of the pupil diameter respectively



HCM overview



Mechanical design



Preliminary system budgets confirm that contrast requirement is achieved